MEMBER BENEFITS

FREE MAPS JOURNAL ACCESS

Enjoy a complimentary online subscription to Meteoritics & Planetary Science (MaPS) and the joint society journal Elements published six times a year.

RESEARCH & COMMUNITY GRANTS

- RESEARCH GRANTS: Support for research activities by student and early-career members.
- COMMUNITY GRANTS: Support for activities and projects that further the goals of the Meteoritical Society, including education and public engagement activities.

DISCOUNTED MEETING REGISTRATION

Receive reduced member rates for the annual Meteoritical Society Meeting.

STUDENT SUPPORT

Engage in new Society initiatives, participate in committees, and help shape the future of meteoritics.

INVOLVEMENT OPPORTUNITIES

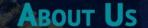
Join committees, participate in Society initiatives, and help shape the future of meteori-

GLOBAL SCIENTIFIC COMMUNITY

The Meteoritical Society actively collaborates with:
The Geochemical Society
The Geological Society of America
The Lunar and Planetary Institute

The Barringer Crater Company

International Union of Geological Sciences Together, we support international research, scientific exchange, and educational outreach in planetary science and meteoritics.



The Meteoritical Society is an international non-profit organization founded in 1933 to advance the study of meteorites and other extraterrestrial materials.

We have over 1,000 members from more than 60 countries, including scientists, students, enthusiasts, collectors, and educators dedicated to understanding the origin and evolution of our Solar System.

Our members investigate meteorites, cosmic dust, asteroids and comets, natural satellites, planets, impact craters, and the origins of the Solar System. Many contribute to sample-return missions, planetary exploration, and public outreach worldwide.

WHAT WE DO

Publish Meteoritics & Planetary Science (MaPS)

Our peer-reviewed journal shares cutting-edge research on meteorites, planetary materials, and Solar System evolution.

Host the Annual Meteoritical Society Meeting (MetSoc)

An international conference bringing together researchers, students, and enthusiasts to share discoveries and build global collaborations.

Promote Global Collaboration

Connecting universities, laboratories, museums, and space agencies such as NASA, JAXA, ESA, and others.

Recognize Excellence

Through Society awards that honor outstanding scientific contributions and community leadership.

Nomenclature Committee

Oversees the official registry, approval, and naming of all recognized meteorites.

Impact Crater Committee

Maintains scientific standards for identifying and evaluating terrestrial impact structures.



NOTE:

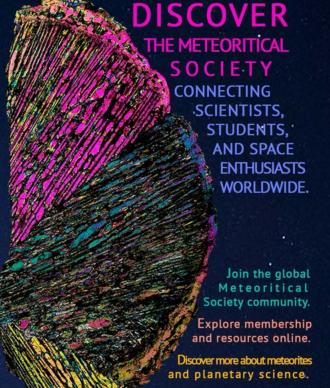
The Meteoritical Bulletin Database is freely accessible to everyone.

JOIN TODAY

The Meteoritical Society metsocsec@meteoritical.org www.MeteoriticalSociety.org



The Meteoritical Society



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THE METEORITICAL SOCIETY www.meteoritical.org



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Asteroid



science since 1933.





anets



WHAT DO METEORITES TELL US? Meteorites are natural archives of the early solar system; their components provide direct evidence of its composition, formation environments, and physical conditions. CAIS - SOLAR SYSTEM'S FIRST SOLIDS Calcium-Aluminum-rich Inclusions (CAIs) are the oldest solids in the solar system. formed near the young Sun. Found in primitive meteorites, they record the early condensation history of the solar nebula. Chondrules – Ancient Molten Droplets • Chondrules are small, spherical particles formed by rapid melting and cooling of dust in the solar nebula. They record high-temperature processes from the early solar system. CHONDRITES - PRIMITIVE METEORITES • Chondrites are undifferentiated stony meteorites made of chondrules, CAIs, matrix, and metal grains. They are divided into carbonaceous (CC) and non-carbonaceous (NC) groups, which formed in distinct regions of the solar system. ACHONDRITES - DIFFERENTIATED METAORITES •
These stony meteorites from differentiated parent bodies experienced melting, crust formation, and igneous activity. Unlike chondrites, they lack chondrules and originate from planets, the Moon, or asteroids with complex geological histories. Pallasites – Stony-Iron Meteorites •-These are meteorites made of metallic iron-nickel and green olivine crystals (peridot). They likely formed at the core-mantle boundary of a differentiated asteroid, offering insight into planetary interiors. RON METEORITES - PLANETARY CORE FRAGMENTS Composed mostly of iron-nickel metal, these meteorites come from the cores of differentiated asteroids. Their structure reveals how

planetary cores form and cool over time.

LUNAR METEORITES .

These Moon rocks are ejected by impacts, found on Earth, and Identified by composition and isotopic signatures matching Apollo samples. They reveal the Moon's crust, volcanism, and impact

MARTIAN METEORITES -

Rocks are blasted off Mars by large impacts and land on Earth. They are identified as Martian by their distinct chemical and isotopic signatures that match spacecraft data. They reveal Mars' geology, volcanism, and past water activity.

NEAR-EARTH ASTEROIDS - SAMPLE RETURN MISSIONS

NEAs are small rocky bodies whose orbits bring them close to Earth. Spacecraft missions like JAXA's Hayabusa2 (Ryugu) and NASA's OSIRIS-REx (Bennu) have returned pristine, uncontaminated samples preserved in their original space environment.

ORGANIC MATTER - COSMIC CARBON

Carbon-based molecules such as amino acids, hydrocarbons, and olecules formed in space before life arose on Earth. Found as soluble or insoluble ma-terial in meteorites, they offer clues to the chemistry that led to life's origins.

ASTEROIDS - ANCIENT ROCKY REMNANTS

They are small, airless, rocky objects that orbit the Sun, mostly found in the asteroid belt between Mars and Jupiter. They are remnants of the early solar system-primitive bodies that never grew large enough to become planets—and range in size from tiny pebbles to hundreds of km across.

COSMIC DUST - TINY GRAINS FROM SPACE

Cosmic dust are microscopic particles of rock and ice drifting through space. These ancient grains, formed in stars and the early solar nebula, continuously fall to Earth, carrying clues to the origins of planets and organic matter.

COMETS - ICY BODIES WITH GLOWING TAILS

Comets are lcy solar system bodies that release gas and dust when near the Sun, forming bright comas and tails. They preserve volatile elements and organic compounds that may have contributed to the emergence of life on Earth.

METEOROIDS - FRAGMENTS TRAVELING THROUGH SPACE

They are solid objects, typically ranging from dust-sized grains to meter-scale fragments, traveling through interplanetary space. Most originate from asteroids or comets and occasionally collide with planets.

FIREBALLS - EXCEPTIONAL BRIGHT METEOR

They are very bright meteors caused by large meteoroids entering Earth's atmosphere. They help scientists track meteorite falls, study en nics, and reconstruct their orbits, revealing where in the solar system the material originated.

METEORS - BRIGHT STREAM

The visible streaks of light ("shooting stars") produced when a meteoroid enters Earth's atmosphere at high velocity causes the surrounding air and the object itself to heat and glow due to frictional ablation.

IMPACT CRATERS - TRACES OF COSMIC IMPACTS

Impact craters form when asteroids or comets collide with a planetary surface at high speed. They preserve the energy and geological effects of these violent events, offering insight into the solar system's dynamic history.

METEO

Fragments of

that survive their intense heating as they pass through Earth's atmosphere and land on the surface are

When freshly fallen, they often have a dark fusion crust—a thin, glassy layer formed by melting during entry. Some are recovered soon after landing (), while others are discovered later on the ground (finds). Meteorites provide direct evidence of the composition and evolution of asteroids and planets.

Asteroid collisions generate powerful impacts that break, melt, and mix surface materials. These events create breccias-meteorites made of rock fragments fused together—which preserve the shock history, sur-

face mixing, and evolution of asteroid.

O White clast in Murchion CM2 chondrite

MPACT CRATER

